

Ladder operators and quasinormal modes in Banados-Teitelboim-Zanelli black holes

Based on 2205.15610 (T.K. and Masashi Kimura)

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Outline

1. Introduction

2. Review: Mass ladder operators

3. Review: Quasinormal modes

4. Quasinormal modes in Banados-Teitelboim-Zanell spacetimes

5. Mass ladder operators in Banados-Teitelboim-Zanell spacetimes

6. Shift of quasinormal mode frequencies

Black hole perturbations

- Test fields in BH spacetimes play an important role in the understanding of phenomena in the strong-gravity regime

e.g., gravitational waves from binary black holes,
time evolution of ultralight bosons around black holes,
linear stabilities of black hole spacetimes,
relaxation phenomena within AdS/CFT, ...

- In many problems, master equations take a form of Schrodinger eq:

$$\left[\frac{d^2}{dx^2} + \omega^2 - V(x) \right] \phi = 0$$

- Mathematical tools in quantum mechanics can be useful in BH perturbation theory
e.g., ladder operators

Ladder operators

- In quantum mechanics,
ladder operators allow to relate the different energy eigenvalues
- Ladder operators in curved spacetimes: **Mass ladder operators** [Cardoso et al, 2017]
[Cardoso et al, 2018]

$$[\square - \mu^2] \Phi = 0 \quad \Rightarrow \quad [\square - (\mu^2 + \delta\mu^2)] D\Phi = 0$$

Mass ladder operator D maps a Klein-Gordon field onto another Klein-Gordon field

This is constructed from spacetime conformal symmetry

Question:

Does a mass ladder operator keep physics determined by boundary conditions?

This work: Application of mass ladder operators to black hole physics

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Review: Mass Ladder operators from spacetime conformal symmetry

- Material: spacetime conformal symmetry

[Wald's textbook]

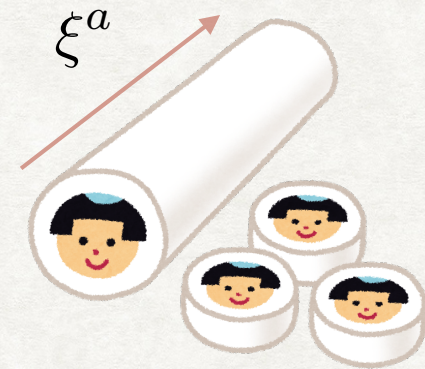
Definition: spacetime symmetry

Spacetime (\mathcal{M}, g_{ab}) possesses symmetry

iff g_{ab} admits an isometry defined by $\varphi_t : \mathcal{M} \rightarrow \mathcal{M}$ such that $\varphi_t^* g_{ab} = g_{ab}$

Isometry group is generated by $x^a \rightarrow \bar{x}^a - \xi^a$

along a **Killing vector field** that satisfies $\mathcal{L}_\xi g_{ab} = 0$

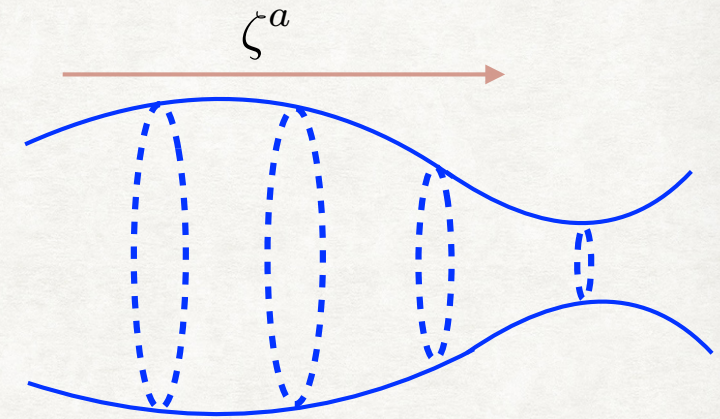


Kintaro-ame

Review: Mass Ladder operators from spacetime conformal symmetry

- What is spacetime conformal symmetry?

Generalization of spacetime symmetry



[Wald's textbook]

Definition: spacetime conformal symmetry

Spacetime (\mathcal{M}, g_{ab}) possesses conformal symmetry

iff g_{ab} admits a conformal isometry defined by $\varphi_t : \mathcal{M} \rightarrow \mathcal{M}$

such that $\varphi_t^* g_{ab} = \exp(2Q) g_{ab}$, where Q is a function on \mathcal{M}

Conformal isometry group is generated by $x^a \rightarrow \bar{x}^a - \zeta^a$

along a **conformal Killing vector field** that satisfies $\mathcal{L}_\zeta g_{ab} = 2Q g_{ab}$

* Conformal Killing vector field is

a Killing vector field of $\tilde{g}_{ab} = \Omega^2 g_{ab}$ with $Q = -2\zeta^a \nabla_a \ln \Omega$

Construction of mass Ladder operators

[Cardoso et al, 2017]

[Cardoso et al, 2018]

• "Closed" condition for conformal Killing vector fields: $\nabla_a \zeta_b = \nabla_b \zeta_a$

• Assumption: $R^a_b \zeta^b = \chi(n-1)\zeta^a$, $\chi \in \mathbb{R}$

e.g., n-dimensional (anti-) de Sitter spacetimes with $\chi = \frac{\Lambda}{n-1}$

• **Mass ladder operator**: $D_k := \mathcal{L}_\zeta - \frac{k}{n} (\nabla_\sigma \zeta^\sigma)$, $k \in \mathbb{R}$

• Commutation relation holds: $[\square, D_k] = \chi(2k+n-2)D_k + 2Q(\square + \chi k(k+n-1))$

$$\Rightarrow D_{k-2} [\square - \mu^2] \Phi = [\square - (\mu^2 + \delta\mu^2)] D_k \Phi, \quad \begin{aligned} \mu^2 &= -\chi k(k+n-1), \\ \delta\mu^2 &= \chi(2k+n-2) \end{aligned}$$

If Φ is a Klein-Gordon field with μ^2 ,

$D_k \Phi$ is also a Klein-Gordon field but with $\mu^2 + \delta\mu^2$

Note: Mass ladder operator is an onto map

Mass Ladder operators

[Cardoso et al, 2017]

[Cardoso et al, 2018]

- Mass ladder operator connects Klein-Gordon fields with different mass squared:

$$[\square - \mu^2] \Phi = 0 \longrightarrow [\square - (\mu^2 + \delta\mu^2)] D_k \Phi = 0$$

$$\mu^2 = -\chi k(k + n - 1),$$

$$\delta\mu^2 = \chi(2k + n - 2)$$

- k is required to be real, so leads to inequalities:

$$\mu^2 \geq \frac{\chi}{4} (n - 1)^2 \quad (\text{for } \chi < 0), \quad \mu^2 \leq \frac{\chi}{4} (n - 1)^2 \quad (\text{for } \chi > 0)$$

Note: In AdS case ($\chi < 0$), the lower bound coincides with the BF bound

- When parametrizing $\mu^2 = -\chi\nu(\nu + n - 1)$ ($\nu \geq -1$),

leads to two solutions, $k_+ = -n + 1 - \nu$, $k_- = \nu$

$$[\square + \chi\nu(\nu + n - 1)] \Phi = 0$$

\nearrow
 \searrow

$$[\square + \chi\tilde{\nu}(\tilde{\nu} + n - 1)] D_{k_+} \Phi = 0$$

$\tilde{\nu} = \nu + 1$ mass raising (lowering) for $\chi < 0$ (> 0)

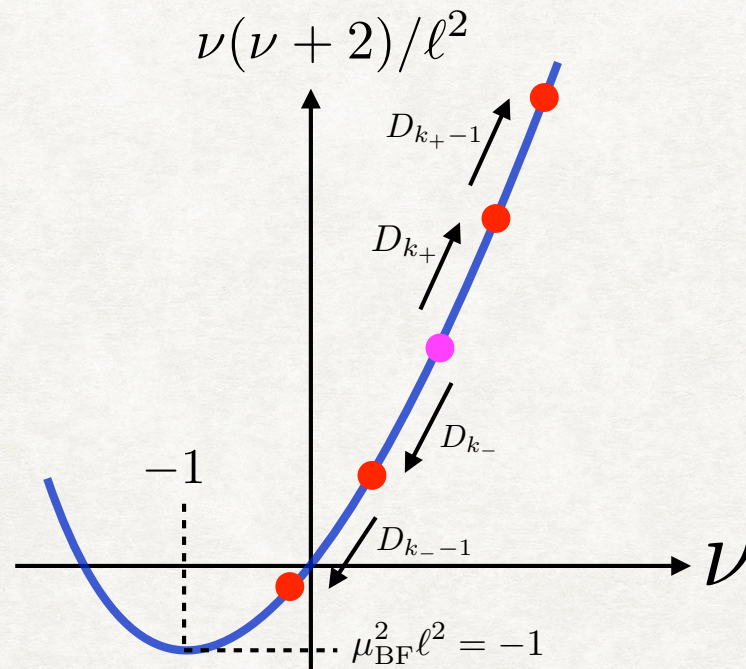
$$[\square + \chi\tilde{\nu}(\tilde{\nu} + n - 1)] D_{k_-} \Phi = 0$$

$\tilde{\nu} = \nu - 1$ mass lowering (raising) for $\chi < 0$ (> 0)

Mass shift

Example: $[\square - \mu^2] \Phi = 0$ on AdS_3 with length scale $\ell := \sqrt{-1/\Lambda}$

Mass parametrization: $\mu^2 = \nu(\nu + 2)/\ell^2$ ($\nu \geq -1$)



Mass ladder operators $D_{k_{\pm}}$ make mass squared raise or lower

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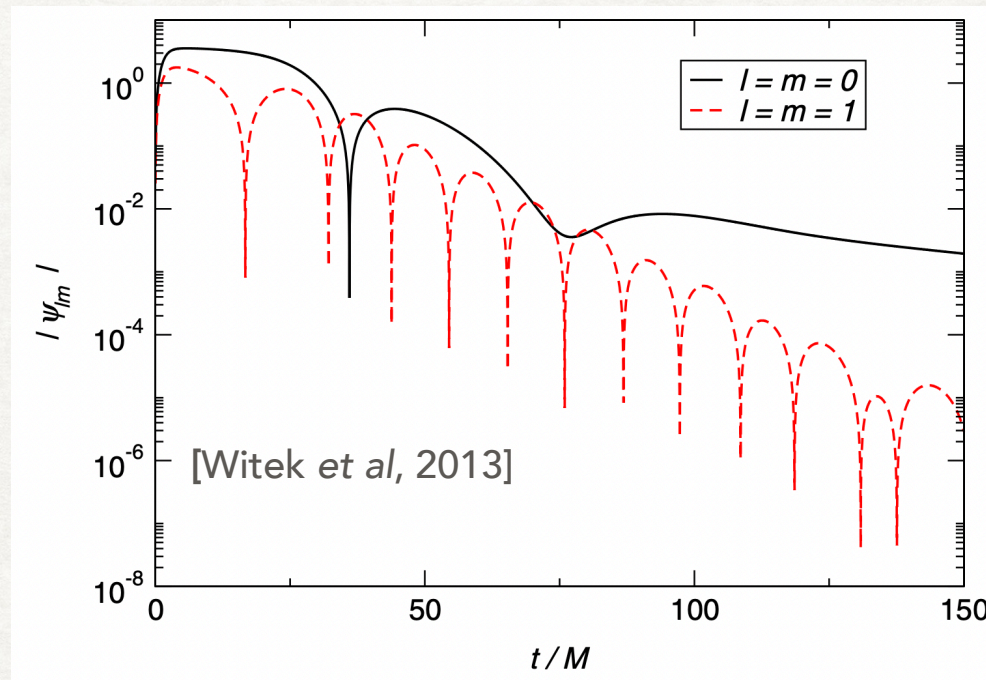
4. Quasinormal modes in Banados-Teitelboim-Zanell spacetimes

5. Mass ladder operators in Banados-Teitelboim-Zanell spacetimes

6. Shift of quasinormal mode frequencies

Quasinormal modes

- Quasinormal modes (QNMs) describe characteristic dynamics of test fields on BH spacetimes as a linear response



- Many applications:

modeling ringdown gravitational waveforms,

[Giesler et al, 2019]

analysis of relaxation phenomena within AdS/CFT,

[Horowitz and Hubeny, 2000]

linear mode stability of BH spacetimes,...

[Regge and Wheeler, 1957]

Brief review: QNMs in asymptotically flat BH spacetimes

- Equation for linear perturbations on asymptotically flat background:

$$\left[\frac{d^2}{dx^2} + \omega^2 - V(x) \right] \phi(x; \omega) = 0$$

e.g., spin- s field perturbation on Schwarzschild backgrounds

$$\left[\frac{d^2}{dx^2} + \omega^2 - V^{(s)} \right] \Phi^{(s)} = 0, \quad \frac{dr}{dx} = 1 - \frac{r_H}{r}$$
$$V^{(s)} = \left(1 - \frac{r_H}{r} \right) \left(\frac{\ell(\ell+1)}{r^2} + \frac{(1-s^2)r_H}{r^3} \right),$$

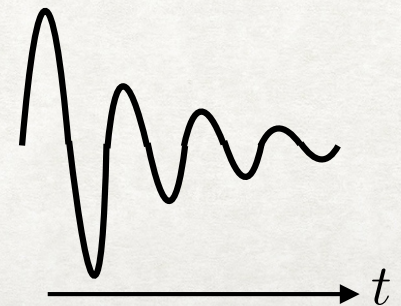
- Appropriate boundary conditions at the horizon $x = -\infty$ and infinity $x = \infty$,

$$\lim_{x \rightarrow -\infty} \phi = e^{-i\omega x} \quad (\text{purely ingoing})$$

$$\lim_{x \rightarrow +\infty} \phi = e^{i\omega x} \quad (\text{purely outgoing})$$

define QNMs and (a discrete set of) QNM frequencies

- QNM frequencies are complex, $\phi \propto e^{-i\text{Re}[\omega]t} e^{\text{Im}[\omega]t}$



Brief review: QNMs as poles of Green's function

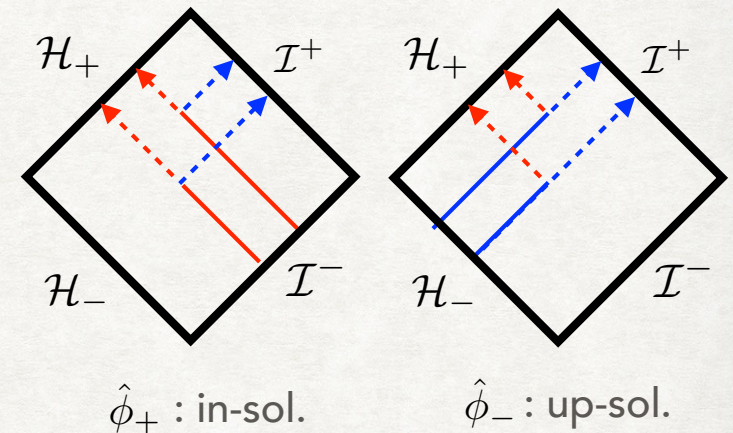
- Equation for linear perturbations in time domain: $[\partial_x^2 - \partial_t^2 - V(x)] \phi(t, x) = 0$

Laplace transform leads to $\left[\frac{d^2}{dx^2} + \omega^2 - V(x) \right] \hat{\phi}(x; \omega) = [i\omega\phi - \partial_t\phi] \Big|_{t=0}$

- Two homogeneous sols. such that

$$\hat{\phi}_+ \simeq \begin{cases} e^{-i\omega t}, & x \rightarrow -\infty, \\ a_1(\omega) e^{-i\omega t} + a_2(\omega) e^{+i\omega t}, & x \rightarrow +\infty \end{cases}$$

$$\hat{\phi}_- \simeq \begin{cases} b_1(\omega) e^{-i\omega t} + b_2(\omega) e^{+i\omega t}, & x \rightarrow -\infty \\ e^{+i\omega t}, & x \rightarrow +\infty, \end{cases}$$



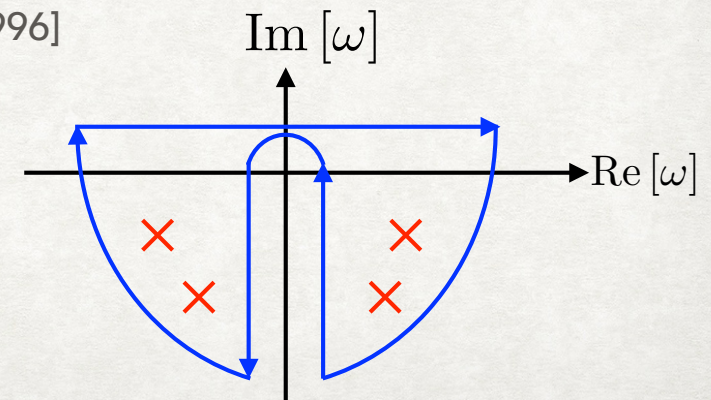
- Green's function in frequency domain: [Leaver,1988]

$$\hat{G}(x, x'; \omega) = \frac{\hat{\phi}_+(x; \omega) \hat{\phi}_-(x'; \omega)}{W(\omega)}$$

[Anderson,1996]

where W is a Wronskian of $\hat{\phi}_+$ and $\hat{\phi}_-$

QNM frequencies are determined by $W = 0$



Brief review: Overtones

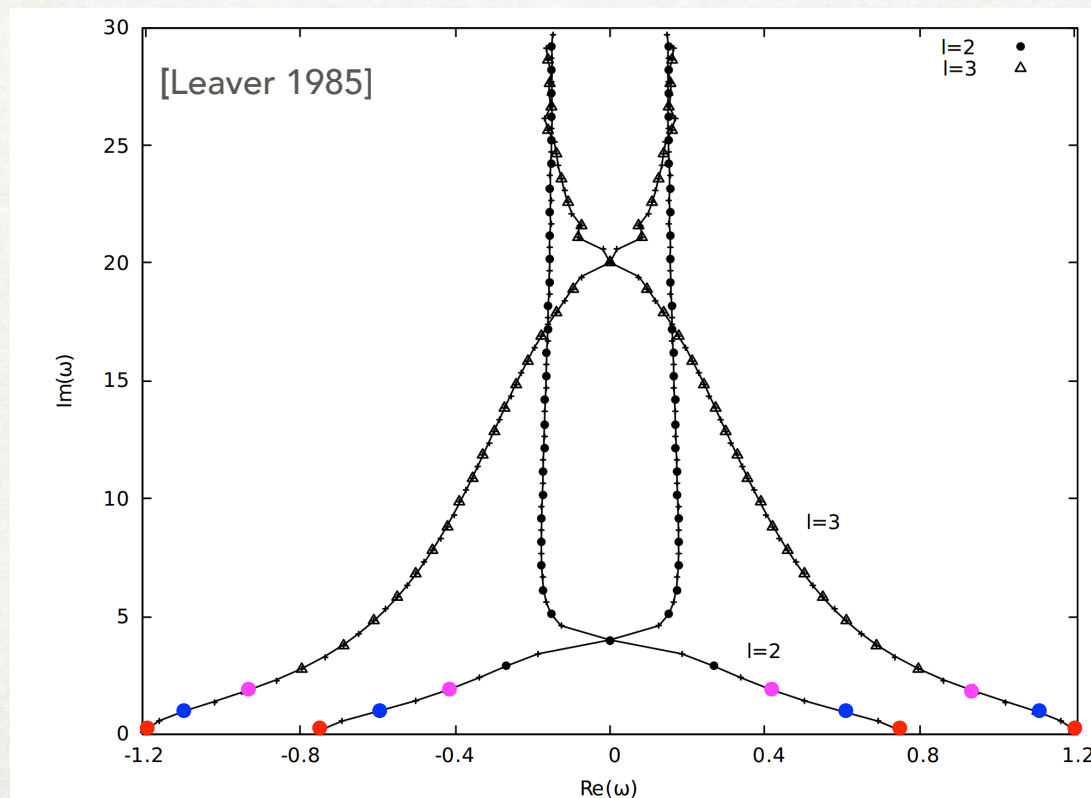
- QNM takes a form:
$$\Phi_{\text{QNM}} = \sum_{\ell=2}^{\infty} \sum_{m=-\ell}^{\ell} \sum_{n=0}^{\infty} \phi_{\ell mn}(x) e^{-i\omega_{\ell mn} t} Y_{\ell m}(\theta, \varphi)$$

For each mode with ℓ ,

there exists a discrete set of modes labeled by $n(= 0, 1, 2, \dots)$: **overtones**

- Index of overtones is defined in the order from the smallest value of $|\text{Im}[\omega]|$

$n = 0$: **fundamental mode**, $n = 1$: **1st overtones**, $n = 2$: **2nd overtones**,...



QNMs in AdS BH spacetimes

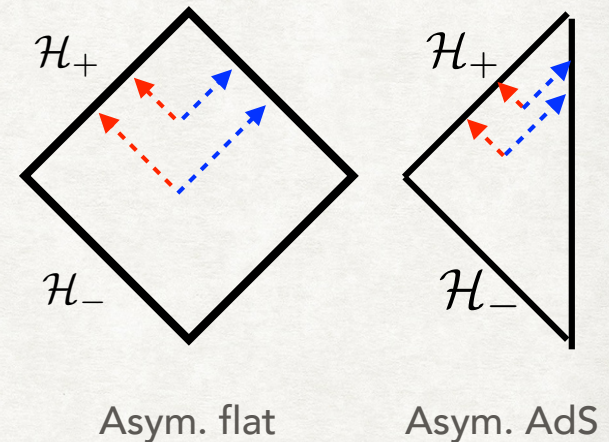
- QNMs in AdS BH spacetimes can be defined in the same manner [Berti *et al*, 2009]
- Variety of boundary condition at infinity exists due to the asymptotic structure

e.g., $\lim_{x \rightarrow \infty} \phi = A + \frac{B}{x^3}$ for massless scalar field

$A = 0$: Dirichlet condition

$B = 0$: Neumann condition

$B = \kappa A, \kappa \in \mathcal{R}$: Robin condition



➡ Rich structure of QNM dynamics appears in AdS (as will be seen later)

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Question and our work

Mass ladder operator works in curved spacetimes with conformal symmetry:

$$[\square - \mu^2] \Phi = 0 \quad \Rightarrow \quad [\square - (\mu^2 + \delta\mu^2)] D\Phi = 0$$

Question:

Does a mass ladder operator keep physics determined by boundary conditions?

We study QNMs of a massive Klein-Gordon field
in Banados-Teitelboim-Zanelli black hole spacetimes

Why BTZ?: BTZ spacetime is the simplest system,
in which QNMs and mass ladder operators can be exactly derived

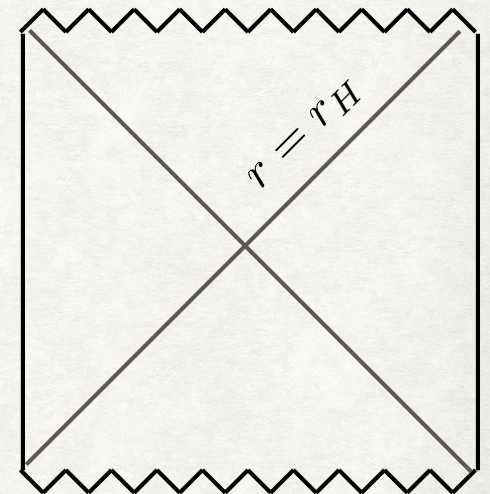
Static Banados-Teitelboim-Zanelli spacetime

- BTZ geometry describes asymptotically AdS black hole spacetimes in 3 dim.
- Line element in (t, r, φ) coordinates: [Banados, Teitelboim, and Zanelli, 1992]

$$ds^2 = -N^2(r) dt^2 + \frac{1}{N^2(r)} dr^2 + r^2 d\varphi^2, \quad N^2(r) = \frac{r^2 - r_H^2}{\ell^2}$$

$$-\infty < t < \infty, \quad r_H < r < \infty, \quad 0 \leq \varphi < 2\pi$$

- Horizon is located at $r = r_H$ such that $N^2(r_H) = 0$
- locally isometric to AdS_3



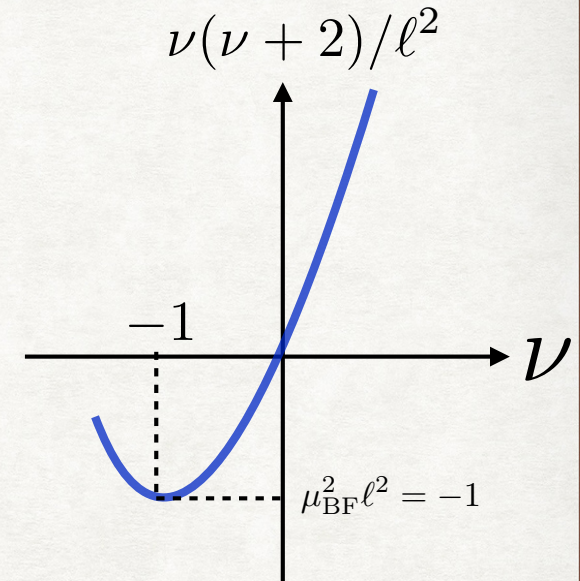
*BTZ BH can rotate but for simplicity we consider the static version

Massive Klein-Gordon fields

- Massive Klein-Gordon field: $\left[\nabla_\mu \nabla^\mu - \frac{\nu(\nu+2)}{\ell^2} \right] \Phi = 0$
 $\nu \geq -1$

- Expanding the field: $\Phi = \sum_m \phi_m(r) e^{-i\omega_m t} e^{im\varphi}$

$$\phi'' + \left(\frac{1}{r} + \frac{(N^2)'}{N^2} \right) \phi' + \frac{1}{N^2} \left(\frac{\omega^2}{N^2} - \frac{m^2}{r^2} - \frac{\nu(\nu+2)}{\ell^2} \right) \phi = 0$$



- Appropriate B.C. selects a discrete set of eigenvalues, i.e., QNM frequencies

At the horizon: Ingoing-wave condition

$$\phi = A \left(1 - \frac{r_H^2}{r^2} \right)^{-\frac{i\omega\ell^2}{2r_H}} \left(\frac{r_H}{r} \right)^{\nu+2} {}_2F_1 \left(a, b; c; 1 - r_H^2/r^2 \right)$$

$$a = \frac{\nu+2}{2} - i \frac{\ell}{2r_H} (\omega\ell - m)$$

$$b = \frac{\nu+2}{2} - i \frac{\ell}{2r_H} (\omega\ell + m)$$

$$c = 1 - i \frac{\omega\ell^2}{r_H}$$

QNMs in BTZ spacetimes

- B.C. at infinity: Dirichlet B.C. for $\nu > -1$ ($\mu^2 > \mu_{\text{BF}}^2$)

$$\phi = A_{\text{I}}(\omega) \left(\frac{r_H}{r}\right)^{-\nu} [1 + \dots] + A_{\text{II}}(\omega) \left(\frac{r_H}{r}\right)^{\nu+2} [1 + \dots]$$

$$a = \frac{\nu+2}{2} - i \frac{\ell}{2r_H} (\omega\ell - m)$$

$$b = \frac{\nu+2}{2} - i \frac{\ell}{2r_H} (\omega\ell + m)$$

$$c = 1 - i \frac{\omega\ell^2}{r_H}$$

$$A_{\text{I}}(\omega) \left(:= A \frac{\Gamma(c)\Gamma(a+b-c)}{\Gamma(a)\Gamma(b)} \right) \text{ vanishes}$$

if $a = -n$ or $b = -n$ for $n = 0, 1, 2, \dots$ due to $1/\Gamma(-n) = 0$

- QNMs frequencies: $\omega_{\text{D}} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 2 + \nu)$ [Cardoso and Lemos, 2001]

$$\phi = A \left(1 - \frac{r_H^2}{r^2}\right)^{-i \frac{\omega_{\text{D}} \ell^2}{2r_H}} \left(\frac{r_H}{r}\right)^{2+\nu} \sum_{k=0}^n \frac{(a)_k (b)_k}{k! (c)_k} \left(1 - \frac{r_H^2}{r^2}\right)^k \text{ where } (\xi)_k \equiv \Gamma(z+k)/\Gamma(z)$$

Imaginary parts are negative, indicating linear mode stability

QNMs in BTZ spacetimes: Other boundary condition

- B.C. at infinity: Neumann B.C. for $\nu > -1$ ($\mu^2 > \mu_{\text{BF}}^2$)

$$\phi = A_{\text{I}}(\omega) \left(\frac{r_H}{r}\right)^{-\nu} [1 + \dots] + A_{\text{II}}(\omega) \left(\frac{r_H}{r}\right)^{\nu+2} [1 + \dots]$$

⇒ $\omega_{\text{N}} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n - \nu)$

Imaginary part can be nonnegative if $\nu \geq 0$ ($\mu^2 \geq 0$),
indicating linear mode instability
due to the presence of non-normalizable mode

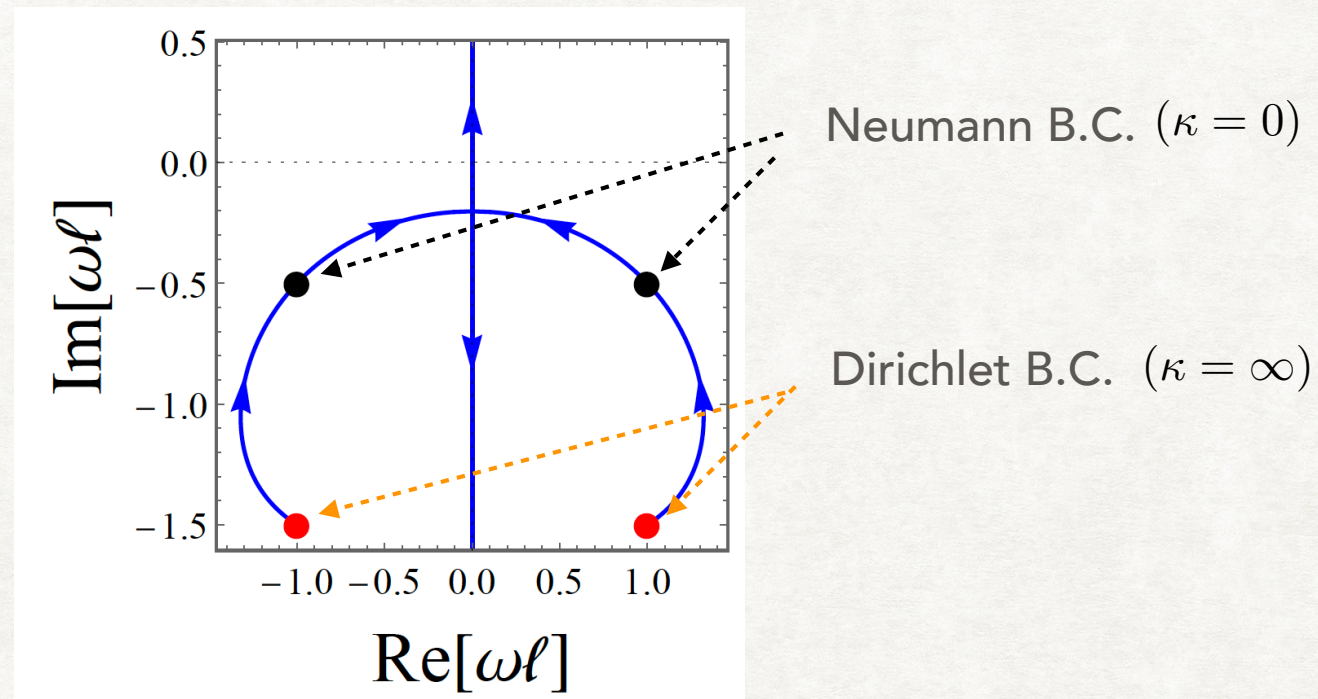
- B.C. at infinity: Robin B.C. for $-1 < \nu < 0$ ($\mu_{\text{BF}}^2 < \mu^2 < 0$)

$$A_{\text{II}}/A_{\text{I}} = \kappa \quad (\kappa \in \mathbb{R})$$

including the Dirichlet $\kappa = \infty$ ($A_{\text{I}} = 0$) and Neumann B.C. $\kappa = 0$ ($A_{\text{II}} = 0$)

In this sense, Robin B.C. is more general and admits rich structure

Fundamental modes in Robin condition



As κ decreases, the trajectories approach the imaginary axis, eventually intersect, and split into two parts

[Ishibashi and Wald, 2004]

There exist growing modes, indicating linear mode instability due to the boundary condition

[TK and Harada, 2021]

QNMs in BTZ spacetimes: BF bound case

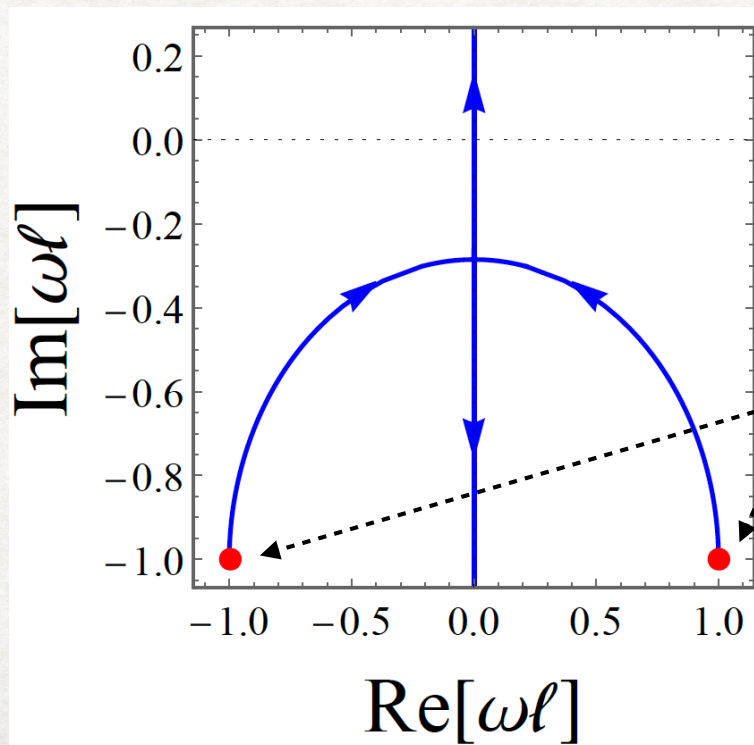
- B.C. at infinity: Dirichlet-Neumann B.C. for $\nu = -1$ ($\mu^2 = \mu_{\text{BF}}^2$) [Ishibashi and Wald, 2004]

$$\phi(r) = A_{\text{I,BF}} \frac{r_H}{r} + A_{\text{II,BF}} \frac{r_H}{r} \ln\left(\frac{r_H}{r}\right) + \mathcal{O}(1/r^3),$$

$$\Rightarrow \omega_{\text{DN}} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 1)$$

- B.C. at infinity: Robin B.C. for $\nu = -1$ ($\mu^2 = \mu_{\text{BF}}^2$)

$$A_{\text{II,BF}}/A_{\text{I,BF}} = 1/\kappa_{\text{BF}} \quad (\kappa_{\text{BF}} \in \mathbb{R})$$



Dirichlet - Neumann B.C. ($\kappa_{\text{BF}} \rightarrow -\infty$)

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Mass ladder operators in BTZ spacetimes

- Mass ladder operators: $D_{i,k} := \mathcal{L}_{\zeta_i} - \frac{k}{3} \nabla_{\mu} \zeta_i^{\mu}$, $k \in \mathbb{R}$

- Four independent CCKVs ($i = 0, 1, 2, 3$) exist in the BTZ spacetime, thus:

$$D_{0,k} = e^{\frac{r_H}{\ell^2} t} \left(\frac{1}{\sqrt{r^2 - r_H^2}} \partial_t - \frac{r\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \partial_r + k \frac{\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \right),$$

$$D_{1,k} = e^{-\frac{r_H}{\ell^2} t} \left(\frac{1}{\sqrt{r^2 - r_H^2}} \partial_t + \frac{r\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \partial_r - k \frac{\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \right),$$

$$D_{2,k} = e^{\frac{r_H}{\ell} \varphi} \left(\frac{r^2 - r_H^2}{\ell r_H} \partial_r + \frac{1}{r} \partial_{\varphi} - k \frac{r}{\ell r_H} \right),$$

$$D_{3,k} = e^{-\frac{r_H}{\ell} \varphi} \left(-\frac{r^2 - r_H^2}{\ell r_H} \partial_r + \frac{1}{r} \partial_{\varphi} + k \frac{r}{\ell r_H} \right).$$

Acting on $\Phi \propto e^{-i\omega t} e^{im\varphi}$, the factors $e^{\pm \frac{r_H^2}{\ell^2} t}$ shift $\omega \rightarrow \omega \pm ir_H/\ell^2$

while $e^{\pm \frac{r_H}{\ell} \varphi}$ break the periodicity to φ



We mainly focus on $D_{0,k}, D_{1,k}$ ($D_{2,k}, D_{3,k}$ still work by multiple action)

Mass ladder operators in BTZ spacetimes

- Commutation relation: $[\nabla_\mu \nabla^\mu, D_{i,k}] = -\frac{2k+1}{\ell^2} D_{i,k} + \frac{2}{3} (\nabla_\mu \zeta_i^\mu) \left[\nabla_\mu \nabla^\mu - \frac{k(k+2)}{\ell^2} \right],$

When choosing $k = k_+ := -2 - \nu$

$$D_{i,k_+-2} \left[\nabla_\mu \nabla^\mu - \frac{\nu(\nu+2)}{\ell^2} \right] \Phi = \left[\nabla_\mu \nabla^\mu - \frac{(\nu+1)(\nu+3)}{\ell^2} \right] D_{i,k_+} \Phi.$$

When choosing $k = k_- := \nu$

$$D_{i,k_--2} \left[\nabla_\mu \nabla^\mu - \frac{\nu(\nu+2)}{\ell^2} \right] \Phi = \left[\nabla_\mu \nabla^\mu - \frac{(\nu-1)(\nu+1)}{\ell^2} \right] D_{i,k_-} \Phi.$$

➡ For the massive Klein-Gordon field Φ with $\nu(\nu+2)/\ell^2$
 $D_{i,k_\pm} \Phi$ is also that with $\tilde{\nu}(\tilde{\nu}+2)/\ell^2$ ($\tilde{\nu} = \nu \pm 1$)

Mass ladder operators and QNM boundary conditions

- QNM with Dirichlet B.C.: $\Phi = A \left(1 - \frac{r_H^2}{r^2}\right)^{-i\frac{\ell^2}{2r_H}\omega_D} \left(\frac{r_H}{r}\right)^{2+\nu} {}_2F_1(a, b; c; 1 - r_H^2/r^2) e^{-i\omega_D t + im\varphi}$,

$$\omega_D = \pm \frac{m}{\ell} - i\frac{r_H}{\ell^2} (2n + 2 + \nu)$$

$$\Phi|_{r \simeq r_H} = 2^{-i\frac{\ell^2}{2r_H}\omega_D} A \left(\frac{r - r_H}{r_H}\right)^{-i\frac{\ell^2}{2r_H}\omega_D} [1 + \mathcal{O}(r - r_H)] e^{-i\omega_D t + im\varphi}. \quad (\text{Ingoing wave})$$

- Acting $D_{0,k_{\pm}}, D_{1,k_{\pm}}$ on the QNMs, at $r \rightarrow r_H$

$$D_{0,k_{\pm}} \Phi = c_{0,k_{\pm}} \left(\frac{r - r_H}{r_H}\right)^{-i\frac{\ell^2}{2r_H}(\omega + i\frac{r_H}{\ell^2})} [1 + \mathcal{O}(r - r_H)] e^{-i(\omega + i\frac{r_H}{\ell^2})t + im\varphi},$$

$$D_{1,k_{\pm}} \Phi = c_{1,k_{\pm}} \left(\frac{r - r_H}{r_H}\right)^{-i\frac{\ell^2}{2r_H}(\omega - i\frac{r_H}{\ell^2})} [1 + \mathcal{O}(r - r_H)] e^{-i(\omega - i\frac{r_H}{\ell^2})t + im\varphi}.$$

Mass ladder operators keep the ingoing-wave condition

Mass ladder operators and QNM boundary conditions

$$\Phi|_{r \simeq \infty} = A_{\text{II}} \left(\frac{r_H}{r} \right)^{2+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i\omega t + im\varphi} \quad (\text{Dirichlet B.C.})$$

- Acting D_{0,k_+}, D_{1,k_+} on the QNMs, at $r \rightarrow \infty$

$$D_{0,k_+} \Phi = c_{0,k_+}^{(\text{D})} \left(\frac{r_H}{r} \right)^{3+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega + i\frac{r_H}{\ell^2})t + im\varphi},$$
$$D_{1,k_+} \Phi = c_{1,k_+}^{(\text{D})} \left(\frac{r_H}{r} \right)^{3+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega - i\frac{r_H}{\ell^2})t + im\varphi}.$$

Asymptotic behaviors of the Klein-Gordon field with $\tilde{\nu}(\tilde{\nu} + 2)/\ell^2$ ($\tilde{\nu} = \nu + 1$)

Mass ladder operators D_{0,k_+}, D_{1,k_+} keep the Dirichlet B.C.

Mass ladder operators and QNM boundary conditions

$$\Phi|_{r \simeq \infty} = A_{\text{II}} \left(\frac{r_H}{r} \right)^{2+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i\omega t + im\varphi} \quad (\text{Dirichlet B.C.})$$

- Acting D_{0,k_-}, D_{1,k_-} on the QNMs, at $r \rightarrow \infty$

$$D_{0,k_-} \Phi = c_{0,k_-}^{(\text{D})} \left(\frac{r_H}{r} \right)^{1+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega + i\frac{r_H}{\ell^2})t + im\varphi},$$

$$D_{1,k_-} \Phi = c_{1,k_-}^{(\text{D})} \left(\frac{r_H}{r} \right)^{1+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega - i\frac{r_H}{\ell^2})t + im\varphi}.$$

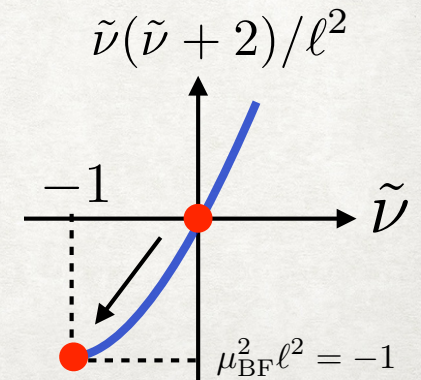
- For $\nu > 0$ ($\mu^2 > 0$) :

Asymptotic behaviors of the Klein-Gordon field with $\tilde{\nu}(\tilde{\nu} + 2)/\ell^2$ ($\tilde{\nu} = \nu - 1$)

Mass ladder operators D_{0,k_-}, D_{1,k_-} keep the Dirichlet B.C.

- For $\nu = 0$ ($\mu^2 = 0$) : $\tilde{\nu} = -1$ corresponds to $\mu_{\text{BF}}^2 \ell^2 = -1$

Dirichlet B.C. changes to the Dirichlet-Neumann B.C.



Mass ladder operators and QNM boundary conditions

$$D_{0,k_-} \Phi = c_{0,k_-}^{(D)} \left(\frac{r_H}{r} \right)^{1+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega + i\frac{r_H}{\ell^2})t + im\varphi},$$

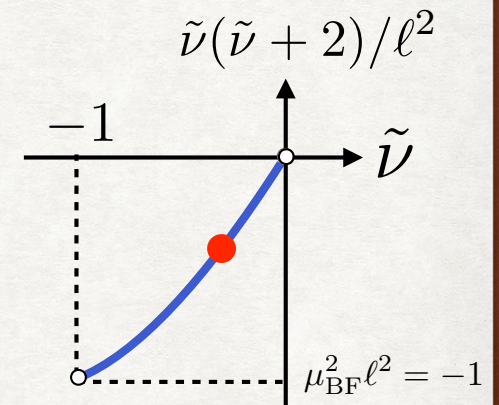
$$D_{1,k_-} \Phi = c_{1,k_-}^{(D)} \left(\frac{r_H}{r} \right)^{1+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega - i\frac{r_H}{\ell^2})t + im\varphi}.$$

For $-1 < \nu < 0$ ($\mu_{\text{BF}}^2 < \mu^2 < 0$) :

The above corresponds to $A_{\text{II}}(\omega) = 0$ (Neumann B.C.)
of the asymptotic behavior

$$\phi = A_{\text{I}}(\omega) \left(\frac{r_H}{r} \right)^{-\tilde{\nu}} [1 + \dots] + \cancel{A_{\text{II}}(\omega) \left(\frac{r_H}{r} \right)^{\tilde{\nu}+2} [1 + \dots]}$$

with $\tilde{\nu} = |\nu| - 1$
($-1 < \tilde{\nu} < 0$)



Dirichlet B.C. changes to Neumann B.C.

Brief summary: Changes of QNM boundary conditions

- At the horizon:

Mass ladder operators $D_{0,k_{\pm}}, D_{1,k_{\pm}}$ keep the ingoing-wave condition

- At infinity:

Mass ladder operators $D_{0,k_{+}}, D_{1,k_{+}}$

Dirichlet B.C. \longrightarrow Dirichlet B.C.

Mass ladder operators $D_{0,k_{-}}, D_{1,k_{-}}$

For $\nu > 0$ ($\mu^2 > 0$) Dirichlet B.C. \longrightarrow Dirichlet B.C.

For $\nu = 0$ ($\mu^2 = 0$) Dirichlet B.C. \longrightarrow Dirichlet-Neumann B.C.

For $-1 < \nu < 0$ ($\mu_{\text{BF}}^2 < \mu^2 < 0$) Dirichlet B.C. \longrightarrow Neumann B.C.

Brief summary: Neumann case

- At infinity:

Mass ladder operators D_{0,k_+}, D_{1,k_+}

Neumann B.C. \longrightarrow Neumann B.C.

Mass ladder operators D_{0,k_-}, D_{1,k_-}

For $\nu > 0$ ($\mu^2 > 0$)

Neumann B.C. \longrightarrow Neumann B.C.

For $\nu = 0$ ($\mu^2 = 0$)

Neumann B.C. \longrightarrow Dirichlet-Neumann B.C.

For $-1 < \nu < 0$ ($\mu_{\text{BF}}^2 < \mu^2 < 0$)

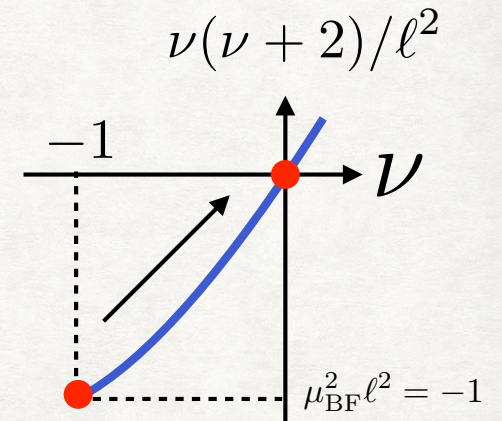
Neumann B.C. \longrightarrow Dirichlet B.C.

Brief summary: Dirichlet-Neumann, Robin cases

- Dirichlet-Neumann case $\nu = -1$ ($\mu^2 = \mu_{\text{BF}}^2$):

Mass ladder operators D_{0,k_+}, D_{1,k_+} ($k_+ = k_-$)

Dirichlet-Neumann B.C. \longrightarrow Dirichlet B.C.



- Robin boundary condition is kept
but the resulting boundary condition parameter is complex

Outline

1. Introduction

2. Review: Mass ladder operators

3. Review: Quasinormal modes

4. Quasinormal modes in Banados-Teitelboim-Zanell spacetimes

5. Mass ladder operators in Banados-Teitelboim-Zanell spacetimes

6. Shift of quasinormal mode frequencies

QNM frequency shift

- Frequency shift from expressions of mass ladder operators

$$D_{0,k} = e^{\frac{r_H}{\ell^2} t} \left(\frac{1}{\sqrt{r^2 - r_H^2}} \partial_t - \frac{r\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \partial_r + k \frac{\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \right),$$

$$D_{1,k} = e^{-\frac{r_H}{\ell^2} t} \left(\frac{1}{\sqrt{r^2 - r_H^2}} \partial_t + \frac{r\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \partial_r - k \frac{\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \right),$$

The factors $e^{\pm \frac{r_H^2}{\ell^2} t}$ suggest $\omega \rightarrow \omega \pm i r_H / \ell^2$

- Example: Shift by D_{0,k_+}

Original QNM frequency with Dirichlet B.C.: $\omega_D = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 2 + \nu)$

Mass parameter shift: $\nu \rightarrow \tilde{\nu} = \nu + 1$

QNM frequency shift: $\omega_D \rightarrow \tilde{\omega}_D = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2(n - 1) + 2 + \tilde{\nu}]$
 $n \rightarrow n - 1$

Note: no "negative overtones" are generated, $D_{0,k_+}[\text{fundamental mode}] = 0$

QNM frequency shift

Operators	D_{0,k_+}	D_{0,k_-}	D_{1,k_+}	D_{1,k_-}
Frequencies	$\omega_D(\nu + 1, n - 1)$	$\omega_D(\nu - 1, n) (\nu > 0)$ $\omega_{DN}(n) (\nu = 0)$ $\omega_N(\nu - 1, n) (-1 < \nu < 0)$	$\omega_D(\nu + 1, n)$	$\omega_D(\nu - 1, n + 1) (\nu > 0)$ $\omega_{DN}(n + 1) (\nu = 0)$ $\omega_N(\nu - 1, n + 1) (-1 < \nu < 0)$

where $\omega_D(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 2 + \nu)$, (Dirichlet B.C.)

$\omega_N(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n - \nu)$, (Neumann B.C.)

$\omega_{DN}(n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 1)$, (Dirichlet-Neumann B.C.)

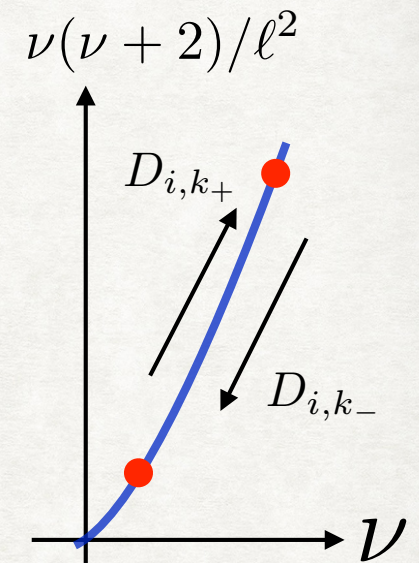
Mass ladder operators change not only mass squared but also indices of overtones

Overtone shift by multiple actions

- Multiple actions $D_{0,k_+ - 1} D_{0,k_-}$ or $D_{1,k_+ - 1} D_{1,k_-}$ keep mass squared but QNM frequencies are shifted:

$$\tilde{\omega} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2(n-1) + 2 + \nu] \quad \text{for } D_{0,k_+ - 1} D_{0,k_-}$$

$$\tilde{\omega} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2(n+1) + 2 + \nu] \quad \text{for } D_{1,k_+ - 1} D_{1,k_-}$$



Note: no "negative overtones" are generated from the fundamental mode

All overtones can be generated from the fundamental mode

Regular solutions generated by multiple actions of $D_{2,k_{\pm}}$, $D_{3,k_{\pm}}$

$$D_{2,k} = e^{\frac{r_H}{\ell} \varphi} \left(\frac{r^2 - r_H^2}{\ell r_H} \partial_r + \frac{1}{r} \partial_{\varphi} - k \frac{r}{\ell r_H} \right),$$

$$D_{3,k} = e^{-\frac{r_H}{\ell} \varphi} \left(-\frac{r^2 - r_H^2}{\ell r_H} \partial_r + \frac{1}{r} \partial_{\varphi} + k \frac{r}{\ell r_H} \right).$$

Factors $e^{\pm \frac{r_H}{\ell} \varphi}$ break the periodicity to φ ;

thus, the single action of $D_{2,k_{\pm}}$, $D_{3,k_{\pm}}$ fails to generate a regular solution

- Multiple actions can remove those singular factors,

e.g., $D_{2,k_+ - 1} D_{3,k_+}$

$$\tilde{\omega} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2(n-1) + 2 + (\nu + 2)] \quad \text{for Dirichlet B.C.}$$

$$\tilde{\omega} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2(n+1) - (\nu + 2)] \quad \text{for Neumann B.C.}$$

e.g., $D_{2,k_- - 1} D_{3,k_+}$

QNM frequency does not change

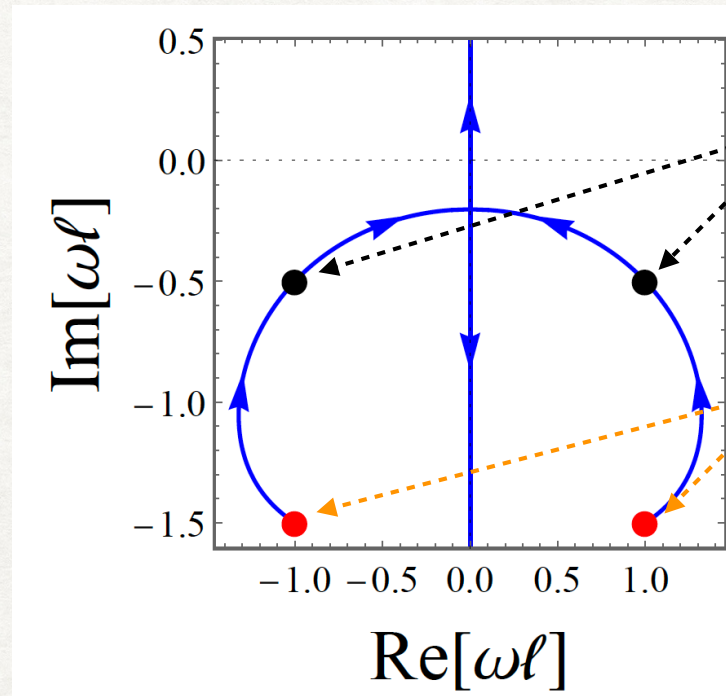
In fact, $D_{2,k_- - 1} D_{3,k_+} = \mathcal{L}_{\xi}$ is a symmetry operator

Other boundary condition: Robin case

[Ishibashi and Wald, 2004]

- Robin B.C. ($-1 < \nu < 0$): $A_{II}/A_I = \kappa$ ($\kappa \in \mathbb{R}$),

$$\phi \simeq A_I(\omega) \left(\frac{r_H}{r}\right)^{-\nu} + A_{II}(\omega) \left(\frac{r_H}{r}\right)^{\nu+2}$$



Neumann B.C. ($\kappa = 0$)

$$\omega_N(0) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [-\nu]$$

Dirichlet B.C. ($\kappa = \infty$)

$$\omega_D(0) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2 + \nu]$$

- Acting the mass ladder operators, $\omega \rightarrow \tilde{\omega} = \omega \pm ir_H/\ell^2$

$$\kappa \rightarrow \tilde{\kappa} \text{ (complex value)}$$

At least, for κ so that ω is purely imaginary,
 $\tilde{\kappa}$ is real, and $\tilde{\omega}$ is also purely imaginary

Summary

We have studied QNMs of massive Klein-Gordon fields in static BTZ spacetimes in terms of mass ladder operators

- Ladder operator is a useful tool to understand mathematical properties of QNMs
- Mass ladder operator change not only mass squared but also QNM frequencies
- In particular, an index of overtones are shifted
- All overtones can be generated from fundamental modes by their multiple actions

Other boundary condition: Neumann case

Operators	D_{0,k_+}	D_{0,k_-}	D_{1,k_+}	D_{1,k_-}
Frequencies	$\omega_N(\nu + 1, n)$	$\omega_N(\nu - 1, n - 1) (\nu > 0)$ $\omega_{DN}(n - 1) (\nu = 0)$ $\omega_D(\nu - 1, n - 1) (-1 < \nu < 0)$	$\omega_N(\nu + 1, n + 1)$	$\omega_N(\nu - 1, n) (\nu > 0)$ $\omega_{DN}(n) (\nu = 0)$ $\omega_D(\nu - 1, n) (-1 < \nu < 0)$

where $\omega_D(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 2 + \nu)$, (Dirichlet B.C.)

$\omega_N(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n - \nu)$, (Neumann B.C.)

$\omega_{DN}(n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 1)$, (Dirichlet-Neumann B.C.)

BF bound case

Operators	$D_{0,-1}$	$D_{1,-1}$	D_0^{BF}
Frequencies	$\omega_{\text{D}}(0, n-1)$	$\omega_{\text{D}}(0, n)$	$\omega_{\text{DN}}(0) \rightarrow \omega_{\text{N}}(0, 0)$

$$\omega_{\text{D}}(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 2 + \nu), \quad (\text{Dirichlet B.C.})$$

$$\omega_{\text{N}}(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n - \nu), \quad (\text{Neumann B.C.})$$

$$\omega_{\text{DN}}(n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 1), \quad (\text{Dirichlet-Neumann B.C.})$$

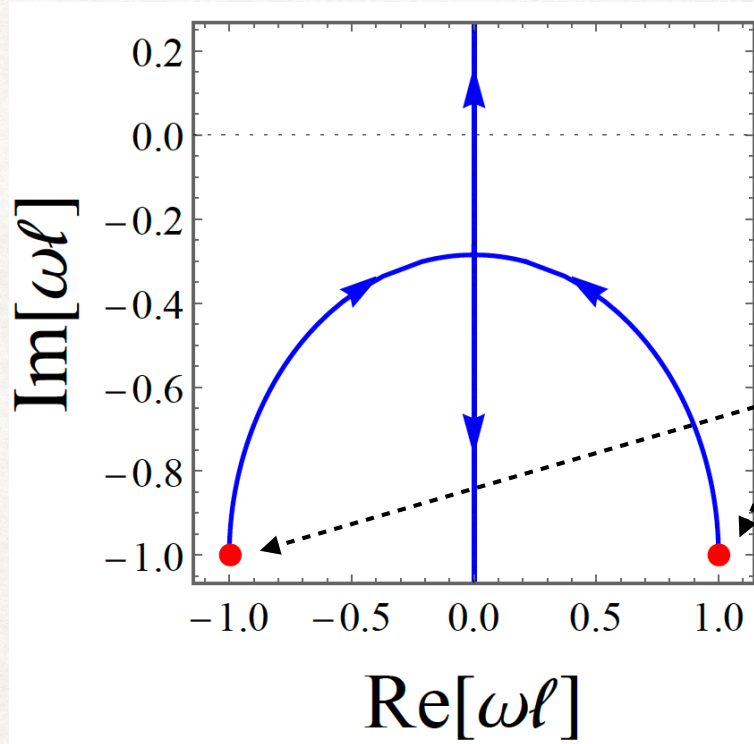
- We find new mass ladder operator only for the fundamental mode in BF bound:

$$D_0^{\text{BF}} \Phi_0 := (\nabla_{\mu} \zeta_0^{\mu}) \Phi_0 = e^{r_H t / \ell^2} \frac{\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \Phi_0,$$

Other boundary condition: Robin case with BF bound

- Robin B.C. ($\nu = -1$): $A_{II,BF}/A_{I,BF} = 1/\kappa_{BF}$ ($\kappa_{BF} \in \mathbb{R}$), [Ishibashi and Wald, 2004]

$$\phi(r) = A_{I,BF} \frac{r_H}{r} + A_{II,BF} \frac{r_H}{r} \ln\left(\frac{r_H}{r}\right) + \mathcal{O}(1/r^3),$$



Dirichlet - Neumann B.C. ($\kappa_{BF} \rightarrow -\infty$)

$$\omega_{\text{DN}}(0) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2}$$

- Acting the mass ladder operators, $\omega \rightarrow \tilde{\omega} = \omega \pm ir_H/\ell^2$
 $\kappa_{BF} \rightarrow \tilde{\kappa}_{BF}$ (complex value)

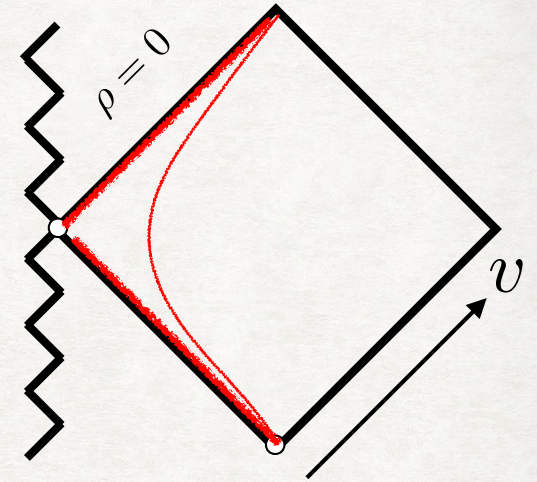
At least, for κ_{BF} so that ω is purely imaginary,
 $\tilde{\kappa}_{BF}$ is real, and $\tilde{\omega}$ is also purely imaginary

Application: scalars in near-horizon geometry

- We have derived conserved quantities along black hole horizons [TK and Kimura 2021]
by exploiting mass ladder operators near the horizon [TK and Kimura 2022]

- Vicinity of black holes with zero Hawking temperature is highly-symmetric geometry called near-horizon geometry:

$$ds^2 = \underbrace{-\lambda_0 \rho^2 dv^2}_{\text{AdS}_2} + 2dv d\rho + \underbrace{\gamma_0 d\Omega_{n-2}^2}_{S^{n-2}}$$



- Reduction of scalars on near-horizon geometry to that on AdS2:

$$\square \Phi = 0 \quad \longrightarrow \quad \left[\square_{\text{AdS}_2} - \frac{\ell(\ell + n - 3)}{\gamma_0} \right] \phi_\ell = 0$$

$$\Phi = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \phi_\ell(v, \rho) Y_{\ell m}(\theta, \varphi)$$

Mass ladder operators connect different multipole modes

- We obtain conservation laws along the horizon: $\partial_v [\partial_\rho D_1 D_2 \cdots D_\ell \phi_\ell] |_{\rho=0} = 0$